

# Chloryl Nitrate: A Novel Product of the OCIO + NO<sub>3</sub> + M Recombination

Randall R. Friedl,\* Stanley P. Sander,

Jet Propulsion Laboratory, California Institute of Technology, Pasadena, California 91109

and Yuk L. Yung

Division of Geological and Planetary Sciences, California Institute of Technology, Pasadena, California 91125  
(Received: May 20, 1992; In Final Form: July 16, 1992)

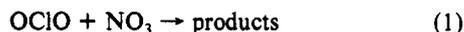
The products of the reaction of OCIO with NO<sub>3</sub> were investigated between 220 and 298 K using a flow reactor and infrared, visible, and ultraviolet analysis. At temperatures below 250 K new infrared and ultraviolet absorption features were observed and assigned to the novel compound chloryl nitrate (O<sub>2</sub>ClONO<sub>2</sub>). Additionally, ClO and NO<sub>2</sub> were observed as reaction products, indicating the existence of a second reaction channel. O<sub>2</sub>ClONO<sub>2</sub> formation predominates at temperatures below 230 K. The reaction rate constant at 220 K is estimated to be on the order of 10<sup>-14</sup> cm<sup>3</sup> molecule<sup>-1</sup> s<sup>-1</sup> in 1-5 Torr of helium. These observations suggest that O<sub>2</sub>ClONO<sub>2</sub> may exist in the terrestrial stratosphere.

## Introduction

The discovery of dramatically elevated concentrations of ClO and OCIO in the polar stratosphere has focused increased attention on the chemistry of the oxides of chlorine.<sup>1-3</sup> Formation of higher oxides has been actively considered; one higher oxide, Cl<sub>2</sub>O<sub>2</sub>, is believed to play a critical role in chlorine-catalyzed destruction of polar ozone.<sup>4</sup> Other oxides, such as Cl<sub>2</sub>O<sub>3</sub>, Cl<sub>2</sub>O<sub>4</sub>, Cl<sub>2</sub>O<sub>6</sub>, and Cl<sub>2</sub>O<sub>7</sub>, have been postulated to serve as temporary reservoirs of chlorine during the early springtime.<sup>5</sup>

An understanding of the partitioning of atmospheric chlorine between reactive and inactive forms is essential for quantitative modeling of polar ozone depletion. Initiation of perturbed chlorine chemistry in the polar atmosphere is thought to occur when inactive chlorine reservoir species HCl and ClONO<sub>2</sub> are converted to photochemically labile Cl<sub>2</sub> and HOCl by heterogeneous reactions occurring on polar stratospheric clouds. Many mechanistic issues regarding the actual chemistry on polar stratospheric clouds remain unresolved, however. Atmospheric modeling has been further hindered by the lack of observational data on chlorine reservoir species in the stratosphere. In particular, the temporal and spatial behaviors of the atmospheric abundances of HCl and ClONO<sub>2</sub> are poorly characterized by field measurements. The existence of other important chlorine reservoir species cannot, at this time, be excluded by the present observational data base.

We have observed new infrared and ultraviolet spectral features arising from the reaction



below temperatures of 250 K. We assign these features to chloryl nitrate (O<sub>2</sub>ClONO<sub>2</sub>), a compound of potential atmospheric significance whose existence was suggested earlier by Christe et al.<sup>6</sup> In this report we present a characterization of the products of reaction 1 and evidence in support of the assignment of chloryl nitrate.

## Experimental Section

Detection of reactants and products was accomplished by long-path absorption in a 1.5-m-long (15-cm-diameter), temperature-controlled, gaseous flow reactor. NO<sub>3</sub> was generated in an upstream prereactor by the reaction of F atoms with HNO<sub>3</sub> or Cl atoms with ClNO<sub>3</sub>. Atomic species were produced by microwave discharge of diatomic precursors. The procedure for synthesizing and handling HNO<sub>3</sub> and ClNO<sub>3</sub> has been discussed previously.<sup>7</sup> OCIO was prepared by passage of Cl<sub>2</sub> through slightly moist NaClO<sub>2</sub> and trapped at 195 K. Products more volatile than OCIO such as Cl<sub>2</sub> were pumped off. A slow flow of helium carried OCIO (approximately 1 Torr vapor pressure at 195 K) into the flow system through a small tubular injector. Using this procedure, products that are less volatile than OCIO, including water, remain

in the trap. No detectable impurities in the OCIO sample were observed using infrared and ultraviolet absorption.

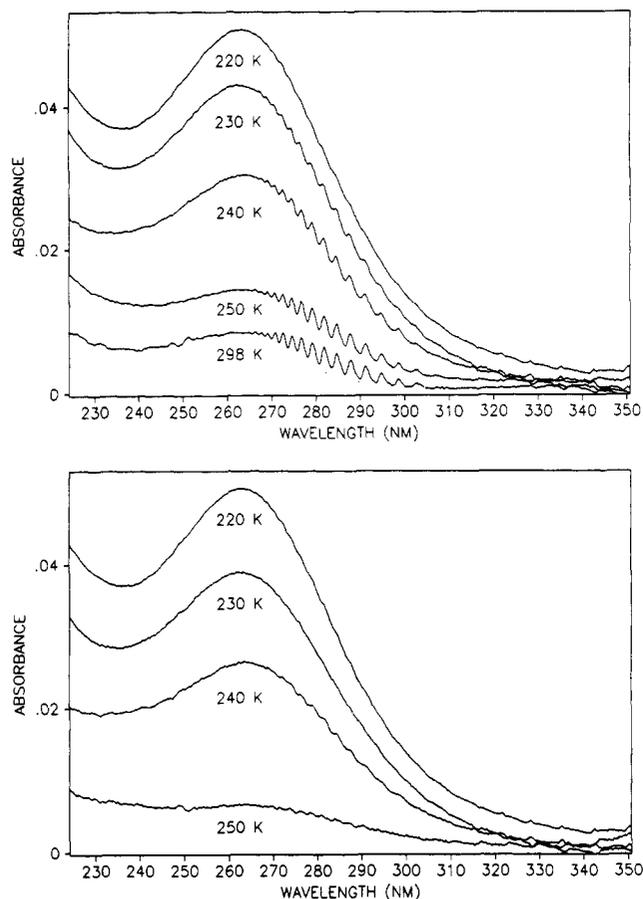
Some additional experiments were performed on the reactions of NO<sub>3</sub> with Cl and ClO. In these experiments ClO was produced by reaction of Cl with Cl<sub>2</sub>O. The Cl<sub>2</sub>O was produced by passage of Cl<sub>2</sub> through HgO.

We conducted experiments between 220 and 298 K and 1 and 8 Torr total pressure of helium buffer gas. The gaseous residence time within the reactor ranged between 1 and 4 s. Total flow rates were between 2 and 6 SLM. Spectroscopic interrogation of the reactor contents was accomplished using a previously described optical arrangement.<sup>8</sup> The infrared optical system consisted of a Bomem DA3+.002 Fourier transform infrared spectrometer coupled to White-type optics in the flow tube (1.5-m base path, 42-m optical path length) and a liquid helium-cooled Cu:Ge detector. The UV/visible optical system consisted of a D<sub>2</sub> or quartz-halogen lamp coupled through the flow tube (1.5-m optical path length) to a diode array spectrometer (PARC 1461/1412 OMA III; 600 grooves/mm grating monochromator) with a 150-nm bandwidth.

The optical systems integrate gaseous absorption along the flow tube axis. The temperature-controlled reactor section encompasses approximately 90% of the optical path. To ensure a uniform reaction temperature, NO<sub>3</sub> was injected at the rear of the reaction cell while OCIO was injected approximately 10 cm downstream, at the beginning of the temperature-controlled region. Integrated absorbances obtained for initial concentrations of NO<sub>3</sub> ranged between 0.02 and 0.3 (as measured at the 662-nm band peak); for OCIO they were between 0.05 and 0.4 (at the 351.3-nm band peak). These absorbances correspond to average flow tube concentrations of (0.6-10) × 10<sup>13</sup> cm<sup>-3</sup> for NO<sub>3</sub> and (2-17) × 10<sup>13</sup> cm<sup>-3</sup> for OCIO.

## Results and Discussion

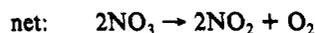
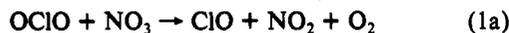
The coaddition of NO<sub>3</sub> and OCIO in the reactor resulted in nearly complete disappearance (>90%) of the visible NO<sub>3</sub> absorption spectrum regardless of the experimental temperature or the ratio of NO<sub>3</sub> and OCIO initial concentrations (varied between 5:1 and 1:5). In contrast, the ultraviolet OCIO absorbance decreased slowly with increasing NO<sub>3</sub> concentration, reaching a minimum value of one-third the initial value (at the highest NO<sub>3</sub>:OCIO ratio). These observations suggest the occurrence of a multiple path or catalytic chain mechanism that effectively destroys NO<sub>3</sub> but conserves OCIO. The kinetics of this process are constrained by the absorbance data obtained for NO<sub>3</sub>. On the basis of the optical data, we estimate that the first-order removal rate of NO<sub>3</sub> was between 1 and 5 s<sup>-1</sup> for the range of OCIO concentrations. The corresponding bimolecular rate coefficient for reaction 1, assuming it to be the rate-limiting step,



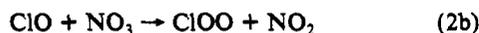
**Figure 1.** Ultraviolet spectra of the products of OCIO + NO<sub>3</sub> taken between 220 and 298 K. (top) Spectra with OCIO absorbance removed. (bottom) Spectra with both ClO and OCIO absorbances removed. The spectrum obtained at 298 K is nearly entirely due to ClO. The remaining spectral features are presumably due to O<sub>2</sub>ClONO<sub>2</sub>, although variations in the spectra below 230 nm may indicate the presence of an additional absorber.

is approximately  $1 \times 10^{-13} \text{ cm}^3 \text{ s}^{-1}$ . Results of a kinetics simulation, based on one possible reaction mechanism, are presented later in this report.

UV spectra of the reaction products are shown in Figure 1. At temperatures near 298 K we observed the appearance of ClO. The production of ClO suggests the following catalytic mechanism:



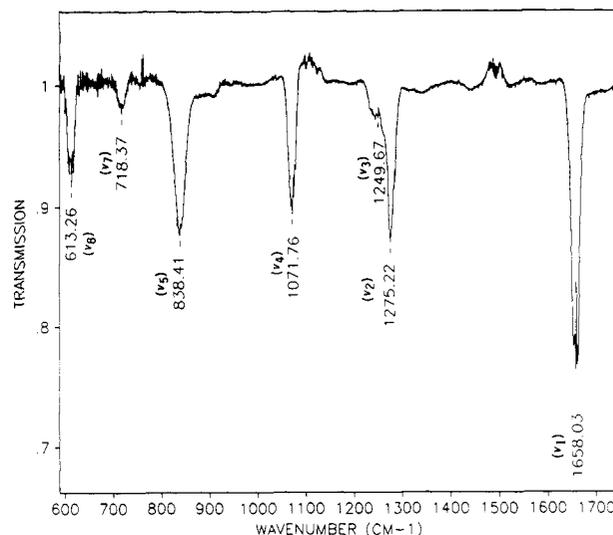
The rate coefficient for reaction 2a is reportedly between  $1 \times 10^{-13}$  and  $4 \times 10^{-13} \text{ cm}^3 \text{ s}^{-1}$ .<sup>9,10</sup> A second channel for reaction 2 has also been proposed,<sup>10</sup> namely



The observed loss of OCIO in the present experiment can be rationalized in terms of the competition between reactions 2a and 2b. The rate of ClOO in our system is dominated by rapid thermal decomposition to yield atomic chlorine and molecular oxygen followed by reaction of the atomic chlorine with either NO<sub>3</sub> or OCIO.



Room temperature infrared spectra of the reaction products were taken at 1-cm<sup>-1</sup> resolution between 600 and 4000 cm<sup>-1</sup>. The choice of instrumental resolution was dictated by signal-to-noise considerations, reactant quantity limitations, and the expectation



**Figure 2.** Infrared difference spectrum of OCIO + NO<sub>3</sub> and NO<sub>3</sub> (before addition of OCIO) taken at 1-cm<sup>-1</sup> resolution and 220 K. The observed bands are attributed to O<sub>2</sub>ClONO<sub>2</sub>. Bands due to OCIO have been removed. The loss of NO<sub>3</sub> is indicated by the positive-going bands at 762 and 1492 cm<sup>-1</sup>.

that an OCIO-NO<sub>3</sub> adduct would have no discernible rovibronic structure. All of the observed bands were attributable to NO<sub>2</sub>, presumably a primary product associated with ClO production, and N<sub>2</sub>O<sub>5</sub>, a secondary product resulting from reaction of NO<sub>2</sub> with NO<sub>3</sub>:



Infrared transitions were not observed for ClO due to the combined effects of small ClO absorption line strengths and reduced instrumental resolution. We conclude that the observed room temperature infrared spectrum is consistent with the reaction scheme 1a, 2a, 2b, and 3-6.

As the reaction temperature was decreased below 298 K, the ultraviolet spectral features due to ClO diminished and a new ultraviolet spectrum was observed with a maximum absorption at 262 nm. The new spectrum is similar to ones observed for the higher chlorine oxides, i.e., Cl<sub>2</sub>O<sub>2</sub>, Cl<sub>2</sub>O<sub>3</sub>, and Cl<sub>2</sub>O<sub>4</sub>. The corresponding infrared spectrum, shown in Figure 2, consisted of seven new bands at 613.3, 718.4, 838.4, 1071.8, 1249.7, 1275.2, and 1658.0 cm<sup>-1</sup>, none of which could be attributed to any of the known higher chlorine oxides.<sup>4,11-15</sup> The relative intensities of the seven bands remained constant for various experimental conditions, indicating the presence of only one absorbing species. We deduced that the absorbing species contains a NO<sub>2</sub> group from the presence of the bands at 1658.0 and 1249.7 cm<sup>-1</sup> and a ClO<sub>2</sub> group from the presence of the bands at 1275.2 and 1071.8 cm<sup>-1</sup>. On the basis of these deductions and our knowledge of the reaction system, we conclude that the unknown product is chloryl nitrate:



Our proposed assignment of the infrared spectrum, assuming C<sub>1</sub> symmetry, is as follows: NO<sub>2</sub> a-stretch (ν<sub>1</sub>) = 1658 cm<sup>-1</sup>, NO<sub>2</sub> s-stretch (ν<sub>3</sub>) = 1250 cm<sup>-1</sup>, ClO<sub>2</sub> a-stretch (ν<sub>2</sub>) = 1275 cm<sup>-1</sup>, ClO<sub>2</sub> s-stretch (ν<sub>4</sub>) = 1072 cm<sup>-1</sup>; ClO stretch (ν<sub>5</sub>) = 838 cm<sup>-1</sup>; NO<sub>2</sub> wag (ν<sub>7</sub>) = 718 cm<sup>-1</sup>; NO stretch (ν<sub>8</sub>) = 613 cm<sup>-1</sup>.

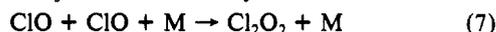
We devised several additional experiments to test our assignment of O<sub>2</sub>ClONO<sub>2</sub>. First, we attempted to verify that reaction 1 was responsible for the production of the observed product. To this end we combined, at 220 K, NO<sub>3</sub> with Cl and ClO and OCIO with NO<sub>2</sub>. None of the new infrared bands ascribed to O<sub>2</sub>ClONO<sub>2</sub> were observed in these trials. No reaction or products were observed for the OCIO + NO<sub>2</sub> reaction. The NO<sub>3</sub> + Cl and NO<sub>3</sub> + ClO reactions produced ClONO<sub>2</sub> (from ClO + NO<sub>2</sub> + M) and trace amounts of OCIO. The relative lack of OCIO product in these experiments indicates that at 220 K the branching ratio of reaction 2 favors ClOO formation, in general agreement with the conclusion of Biggs et al.<sup>10</sup>

TABLE I: Experimental and Modeled Kinetics Results for OCIO + NO<sub>3</sub><sup>a</sup>

<i>T</i> (K)	<i>P</i> (Torr)	rxn time (s)	column abundance (10 <sup>15</sup> molecules cm <sup>-2</sup> )					ClONO <sub>2</sub> abs at 262 nm
			NO <sub>3</sub> initial <sup>b</sup>	NO <sub>3</sub> final <sup>c</sup>	OCIO initial	OCIO final	ClO	
298	3.9	3.2	6.5	1.0 (0.8)	4.7	3.3 (2.6)	1.2 (1.5)	<0.005 (0.04)
			12.0	1.5 (1.3)	4.7	2.4 (2.0)	1.6 (1.7)	<0.005 (0.05)
	3.9	3.2	15.0	2.0 (1.6)	4.7	2.0 (1.8)	2.1 (1.8)	<0.005 (0.06)
	4.3	3.1	12.0	nm <sup>d</sup> (1.2)	10.6	6.1 (6.4)	1.7 (2.6)	<0.005 (0.09)
	4.4	2.9	10.0	1.6 (1.0)	17.3	13.4 (12.4)	1.9 (2.9)	<0.005 (0.12)
220	3.5	4.2	8.7	0.9 (1.3)	5.8	3.1 (2.2)	<0.2 (0.3)	0.032 (1.67)
			13.9	1.3 (1.8)	5.8	1.7 (1.6)	<0.2 (0.4)	0.036 (2.04)
	3.7	3.9	3.5	nm (0.5)	15.5	14.0 (12.6)	<0.2 (0.03)	0.010 (2.11)
	3.7	3.9	12.6	nm (0.8)	16.0	9.3 (7.9)	<0.2 (0.2)	0.032 (3.83)
	4.3	3.5	12.6	nm (1.4)	19.3	14.0 (11.4)	<0.2 (0.2)	0.061 (4.61)

<sup>a</sup>Numbers in parentheses are column abundances (in units of 10<sup>15</sup> molecules cm<sup>-2</sup>) derived from the kinetics simulations using rate coefficients discussed in the text. <sup>b</sup>Denotes column abundance with only one reactant in cell. <sup>c</sup>Denotes column abundance with both reactants in cell. <sup>d</sup>Not measured.

As a second test, we investigated the possibility that Cl<sub>2</sub>O<sub>2</sub> or Cl<sub>2</sub>O<sub>3</sub> formation might account for the observed ultraviolet absorption in the NO<sub>3</sub>-OCIO reaction system:



We added either ClO or ClO + OCIO to the reactor at 220 K and recorded ultraviolet and infrared spectra that are similar to those reported by Burkholder et al.<sup>13</sup> The results of these tests demonstrated that Cl<sub>2</sub>O<sub>2</sub> and Cl<sub>2</sub>O<sub>3</sub> could not have been solely responsible for the ultraviolet absorptions observed in the NO<sub>3</sub> + OCIO system without producing clearly observable infrared bands.

A computer-based kinetics simulation that employed reactions 1-8 was performed on the NO<sub>3</sub> + OCIO reaction system. The purpose of this endeavor was to further establish the plausibility of the deduced reaction mechanism and to derive refined estimates of the reaction rate parameters for reaction 1. The rate constants for reactions 1a and 1b and the branching ratio of reaction 2 were the only variable parameters in this simulation.<sup>9,10,16</sup> A qualitatively good description of the observed NO<sub>3</sub>, OCIO, and ClO behavior at 298 K was obtained with  $k_{1a} = 2 \times 10^{-14}$  cm<sup>3</sup> s<sup>-1</sup>,  $k_{2a} = 1 \times 10^{-13}$  cm<sup>3</sup> s<sup>-1</sup>, and  $k_{2b} = 4 \times 10^{-13}$  cm<sup>3</sup> s<sup>-1</sup>. The observed behavior of ClO as a function of temperature was best described by employing a ratio of  $k_{1b}/k_{1a}$  that increased with decreasing temperature and reached approximately 10 at 220 K. A satisfactory fit to the observed data at 220 K and 4 Torr was obtained with  $k_{1a} = 2 \times 10^{-15}$  cm<sup>3</sup> s<sup>-1</sup> and  $k_{1b} = 2 \times 10^{-14}$  cm<sup>3</sup> s<sup>-1</sup>. A summary of the experimental and modeled kinetics data is shown in Table I. In addition, we estimate that the absorption cross section for O<sub>2</sub>ClONO<sub>2</sub> at 262 nm and 220 K is approximately  $1.3 \times 10^{-17}$  cm<sup>2</sup>, based on the measured absorbances and the predicted column abundances shown in Table I.

The derived kinetics parameters are consistent with a recent report by Becker et al.<sup>9</sup> that places an upper limit of  $7 \times 10^{-14}$  cm<sup>3</sup> s<sup>-1</sup> on  $k_1$  at 298 K. However, we note that the conditions employed by Becker et al. (OCIO decays in excess NO<sub>3</sub>) should be complicated by regeneration of OCIO due to reaction 2a. We are in substantial disagreement with the upper limit value of  $1 \times 10^{-15}$  cm<sup>3</sup> s<sup>-1</sup> reported for  $k_1$  by Biggs et al.<sup>10</sup> No experimental details are given by Biggs et al.; consequently, we are unable to critically compare the two experiments.

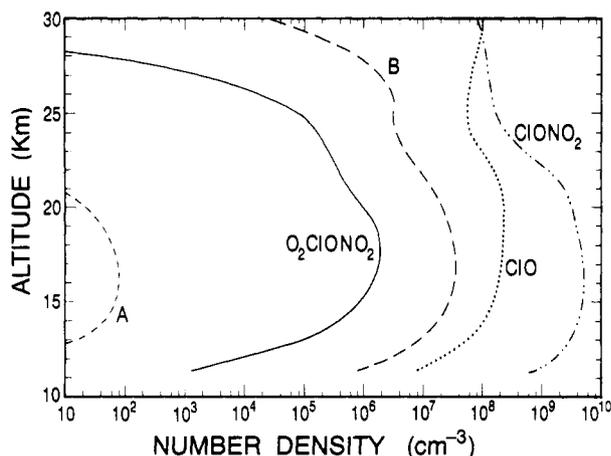
The derived value of  $k_{1b}$  at 220 K was used in conjunction with unimolecular reaction theory<sup>17,18</sup> to estimate the strength of the bond between NO<sub>3</sub> and OCIO in O<sub>2</sub>ClONO<sub>2</sub>. Following the

method used in ref 18, we input the following parameters for the calculation: the seven measured O<sub>2</sub>ClONO<sub>2</sub> vibrational frequencies and eight additional frequencies (740, 530, 480, 355, 280, 190, 115, 100 cm<sup>-1</sup>) derived from a normal-coordinate analysis<sup>19</sup> of O<sub>2</sub>ClONO<sub>2</sub>; an estimated entropy value for O<sub>2</sub>ClONO<sub>2</sub> of 84 eu; an estimated energy difference of 265 cm<sup>-1</sup> between the reaction enthalpy and the critical energy for O<sub>2</sub>ClONO<sub>2</sub> decomposition. The O<sub>2</sub>ClONO<sub>2</sub> bond strength derived from this method is  $18 \pm 3$  kcal mol<sup>-1</sup>. This value is similar to that of ClOOC1 but substantially less than that of ClONO<sub>2</sub>. Consequently, we expect the thermal stability of O<sub>2</sub>ClONO<sub>2</sub> to approach that of ClOOC1.

### Atmospheric Implications

The value derived for  $k_{1b}$  from the kinetics simulations corresponds to a third-order rate constant on the order of  $10^{-31}$  cm<sup>6</sup> s<sup>-2</sup>. This value is consistent with rate parameters for other known reactions of ClO<sub>x</sub> with NO<sub>x</sub> such as ClO + NO<sub>2</sub> and Cl + NO<sub>2</sub>.<sup>16</sup> Accordingly, we believe that the high-pressure limit values of  $k_{1b}$  will approach  $10^{-11}$  cm<sup>3</sup> s<sup>-1</sup>. At temperatures and pressures characteristic of the earth's stratosphere, we expect the value of  $k_{1b}$  to be on the order of  $10^{-12}$  cm<sup>3</sup> s<sup>-1</sup>. In the polar stratosphere this process may serve as a significant sink for active chlorine by converting photolytically labile OCIO into potentially inactive O<sub>2</sub>ClONO<sub>2</sub>. We have modeled the effect of reaction 1b on polar chlorine chemistry using the California Institute of Technology-Jet Propulsion Laboratory one-dimensional photochemical model<sup>20,21</sup> with the exception that the model atmosphere adopted is that of the "collar" region of the polar vortex,<sup>22</sup> where O<sub>2</sub>ClONO<sub>2</sub> is expected to be most abundant. The atmospheric conditions and rate parameters used in the model were similar to those used previously.<sup>5</sup> The rate coefficient for the formation of O<sub>2</sub>ClONO<sub>2</sub> was set equal to that of ClONO<sub>2</sub>. The O<sub>2</sub>ClONO<sub>2</sub> thermal decomposition rate coefficient was calculated using the formation rate coefficient and an estimate of the equilibrium constant for reaction 1b [ $K_{1b} = (1 \times 10^{-28} \exp(9300/T))$ ] that was derived from the unimolecular theory calculation discussed previously. The photodissociation coefficient for O<sub>2</sub>ClONO<sub>2</sub> was set equal to that of HNO<sub>3</sub> since the observed ultraviolet spectrum of O<sub>2</sub>ClONO<sub>2</sub> was unmeasurably small above 300 nm.

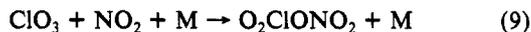
An example of the model results<sup>23</sup> for noontime vertical profiles of O<sub>2</sub>ClONO<sub>2</sub> is shown in Figure 3. There is little variation in the O<sub>2</sub>ClONO<sub>2</sub> profile over a diurnal cycle. The concentrations of the major chlorine species ClO and ClONO<sub>2</sub> in the model are consistent with the column density constraints deduced by Toon et al.<sup>22</sup> in the Antarctic vortex of 1987. A sensitivity study has



**Figure 3.** Model calculations of the vertical profiles of  $\text{O}_2\text{ClONO}_2$ ,  $\text{ClO}$ , and  $\text{ClONO}_2$  in the antarctic stratosphere during noon conditions (second week of spring). Vertical profiles of  $\text{O}_2\text{ClONO}_2$  labeled A and B reflect variations of  $\pm 2 \text{ kcal mol}^{-1}$  in the estimated bond energy (see text).

been carried out on the bond energy of  $\text{O}_2\text{ClONO}_2$ . Curve A (Figure 3) shows the results for  $\text{O}_2\text{ClONO}_2$  if the bond energy is  $2 \text{ kcal mol}^{-1}$  smaller than the standard value. Curve B gives the corresponding results if the bond energy is  $2 \text{ kcal mol}^{-1}$  larger. This large range of uncertainty is the principal source of uncertainty in the modeling of  $\text{O}_2\text{ClONO}_2$ .

The derived column densities of  $\text{O}_2\text{ClONO}_2$  are significantly less than those obtained for  $\text{ClONO}_2$  and  $\text{ClO}$ . Consideration of other reactions which form  $\text{O}_2\text{ClONO}_2$  may increase the atmospheric abundance, however. For example, we propose that the reaction of  $\text{ClO}_3$  with  $\text{NO}_2$  will produce  $\text{O}_2\text{ClONO}_2$ :



Currently we are investigating atmospheric pathways for formation of  $\text{ClO}_3$ .

**Acknowledgment.** Part of the research described in this report was carried out at the Jet Propulsion Laboratory, California

Institute of Technology, under contract to the National Aeronautics and Space Administration. Y.L.Y. was supported by NASA Grant NAGW-413 to the California Institute of Technology.

#### References and Notes

- (1) de Zafra, R. L.; Jaramillo, M.; Parrish, A.; Solomon, P.; Conner, B.; Barrett, J. *Nature* **1987**, *328*, 408.
- (2) Brune, W. H.; Anderson, J. G.; Chan, K. R. *J. Geophys. Res.* **1989**, *94*, 16649.
- (3) Solomon, S.; Mount, G. H.; Sanders, R. W.; Schmeltekopf, A. L. *J. Geophys. Res.* **1987**, *92*, 8329.
- (4) Molina, L. T.; Molina, M. J. *J. Phys. Chem.* **1987**, *91*, 433.
- (5) Sander, S. P.; Friedl, R. R.; Yung, Y. L. *Science* **1989**, *249*, 1095.
- (6) Christe, K. O.; Wilson, W. W.; Wilson, R. P. *Inorg. Chem.* **1989**, *28*, 675.
- (7) Friedl, R. R.; Sander, S. P. *J. Phys. Chem.* **1987**, *91*, 2721.
- (8) Lang, V. I.; Sander, S. P.; Friedl, R. R. *J. Mol. Spectrosc.* **1988**, *132*, 89.
- (9) Becker, E.; Wille, U.; Rahman, M. M.; Schindler, R. N. *Ber. Bunsen-Ges. Phys. Chem.* **1991**, *95*, 1173.
- (10) Biggs, P.; Harwood, M. H.; Paar, A. D.; Wayne, R. P. *J. Phys. Chem.* **1991**, *95*, 7746.
- (11) Schack, C. J.; Pilipovich, D. *Inorg. Chem.* **1970**, *9*, 1387.
- (12) Christe, K. O.; Schack, C. J.; Curtis, E. C. *Inorg. Chem.* **1971**, *10*, 1589.
- (13) Burkholder, J. B.; Orlando, J. J.; Howard, C. J. *J. Phys. Chem.* **1990**, *94*, 687.
- (14) Jansen, M.; Tobias, K. M.; Willner, H. *Naturwissenschaften* **1986**, *73*, 734.
- (15) Witt, J. D.; Hammaker, R. M. *J. Chem. Phys.* **1973**, *58*, 303.
- (16) DeMore, W. B.; Sander, S. P.; Golden, D. M.; Molina, M. J.; Hampson, R. F.; Kurylo, M. J.; Howard, C. J.; Ravishankara, A. R. *JPL Publication 90-1*; Jet Propulsion Laboratory: Pasadena, CA, 1990.
- (17) Troe, J. *J. Phys. Chem.* **1979**, *83*, 114.
- (18) Colussi, A. J. *J. Phys. Chem.* **1990**, *94*, 8922.
- (19) Computer analysis performed using program QCMP067 supplied by the Quantum Chemistry Program Exchange, Indiana University. The input geometrical parameters and force constants for  $\text{O}_2\text{ClONO}_2$  were estimated from  $\text{ClONO}_2$  (Miller, R. H.; Bernitt, D. L.; Hisatsune, I. C. *Spectrochim. Acta* **1967**, *23A*, 223) and  $\text{Cl}_2\text{O}_3$  (author's unpublished data).
- (20) Allen, M.; Yung, Y. L.; Waters, J. *J. Geophys. Res.* **1981**, *86*, 3617.
- (21) Froidevaux, L.; Allen, M.; Yung, Y. L. *J. Geophys. Res.* **1985**, *90*, 12999.
- (22) Toon, G. C.; Farmer, C. B.; Lowes, L. L.; Schaper, P. W.; Blavier, J. F.; Norton, R. H. *J. Geophys. Res.* **1990**, *94*, 16571.
- (23) The maximum concentrations of  $\text{ClO}$ ,  $\text{BrO}$ , and  $\text{NO}_x$  in the model are 1.5 ppbv, 5 pptv, and 10 pptv, respectively. The model results are for the beginning of September.

## Polarized Fourier Transform Infrared Microscopy as a Tool for Structural Analysis of Adsorbates in Molecular Sieves

F. Schüth

*Institut für Anorganische Chemie und Analytische Chemie der Johannes Gutenberg-Universität, Mainz, Becherweg 24, 6500 Mainz, Germany (Received: May 26, 1992; In Final Form: July 28, 1992)*

Using FTIR microscopy with polarized IR radiation on silicalite I single crystals fully loaded with *p*-xylene, the existence of an ordered adsorbate could be proven for the first time by IR spectroscopy. By analyzing the polarized absorption bands the orientation of the *p*-xylene molecules relative to the host structure could be determined. The results agree well with structural data obtained from X-ray diffraction experiments. These first results suggest that polarized IR microscopy could develop into a powerful tool for the analysis of adsorbate structures, assisting in complete structure resolution by diffraction techniques.

### Introduction

Polarized IR spectroscopy is a well-known technique for the analysis of orientated molecules since about 50 years ago.<sup>1</sup> It can be used either for assignment of IR bands to certain modes, if the orientation of molecules with respect to the IR beam is known (see, for instance, the work of Zelei and Dobos on para-substituted benzene derivatives<sup>2-4</sup>), or for structural analysis, if bands have already been assigned unanimously to certain modes, e.g., for orientation in polymers,<sup>5</sup> for analysis of details in the crystal

structure.<sup>6</sup> To our best knowledge, however, this method has until now never been applied to analyze the orientation of adsorbates in molecular sieves. Usually for this purpose X-ray diffraction is used.<sup>7</sup> Occasionally also neutron diffraction<sup>8</sup> or one of these methods in combination with NMR spectroscopy have been applied.<sup>9</sup> Since experiments and data analysis in these experiments can be difficult and time consuming, it would be very advantageous to use a method which allows a relatively simple determination of the orientation of molecules in a crystal. This would exclude